



太陽系有幾顆行星？ (planet formation)

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The Solar System

Planets are almost on the same plane.

Planets orbit the Sun in the same direction.

The Sun is rotating in the same direction as the planets' orbital movement.

Planets form in a disk!

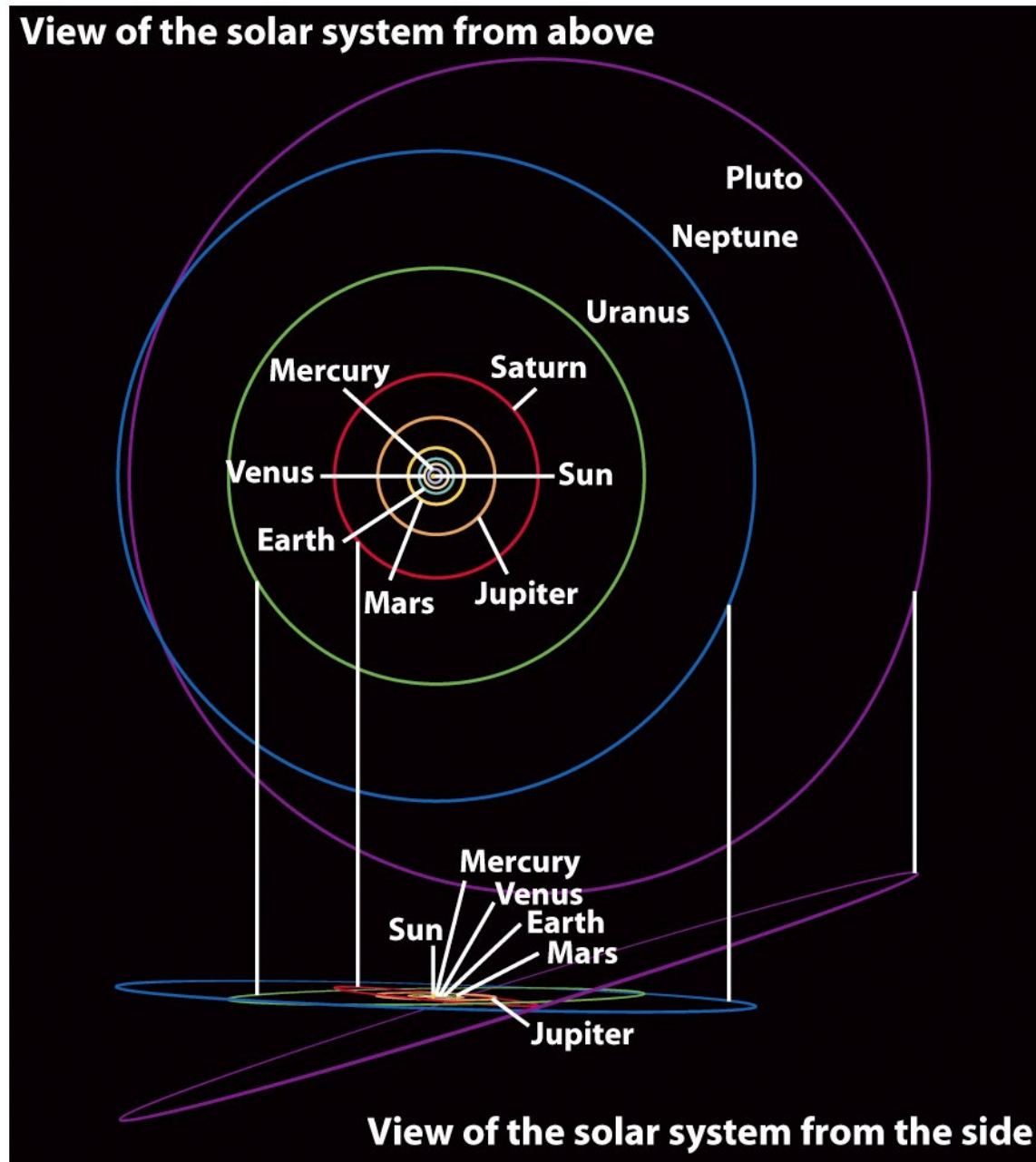
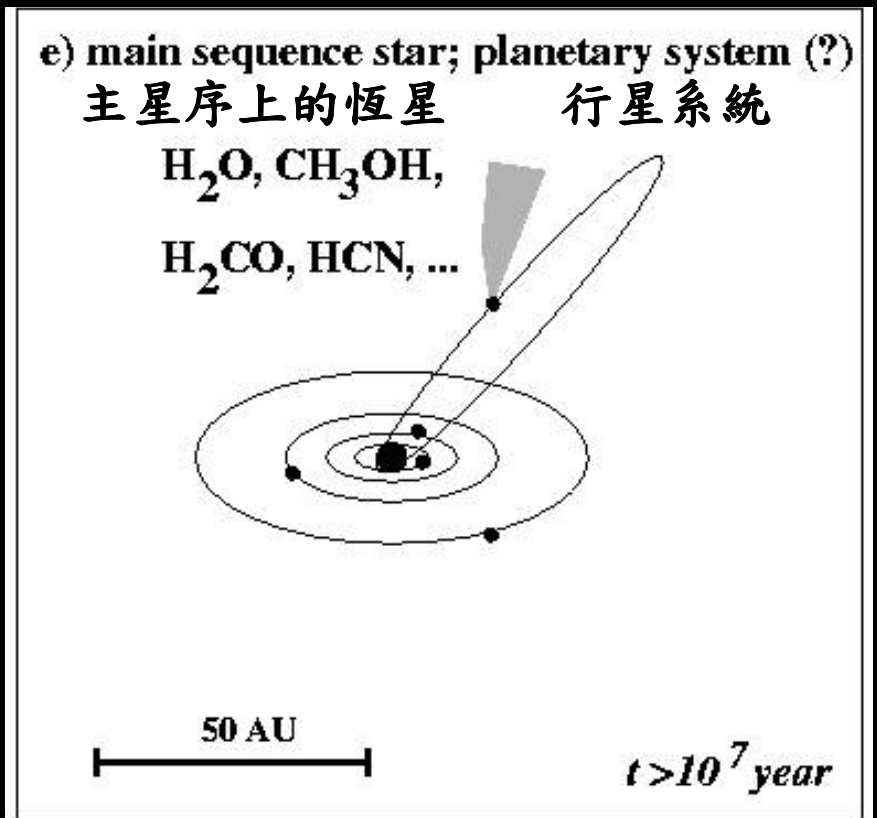
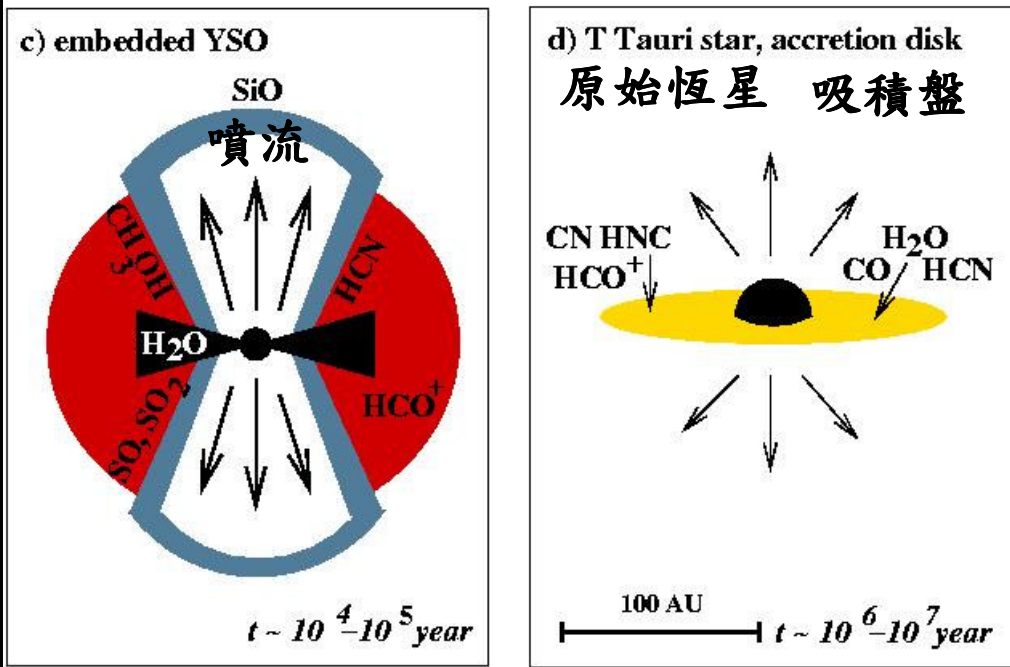
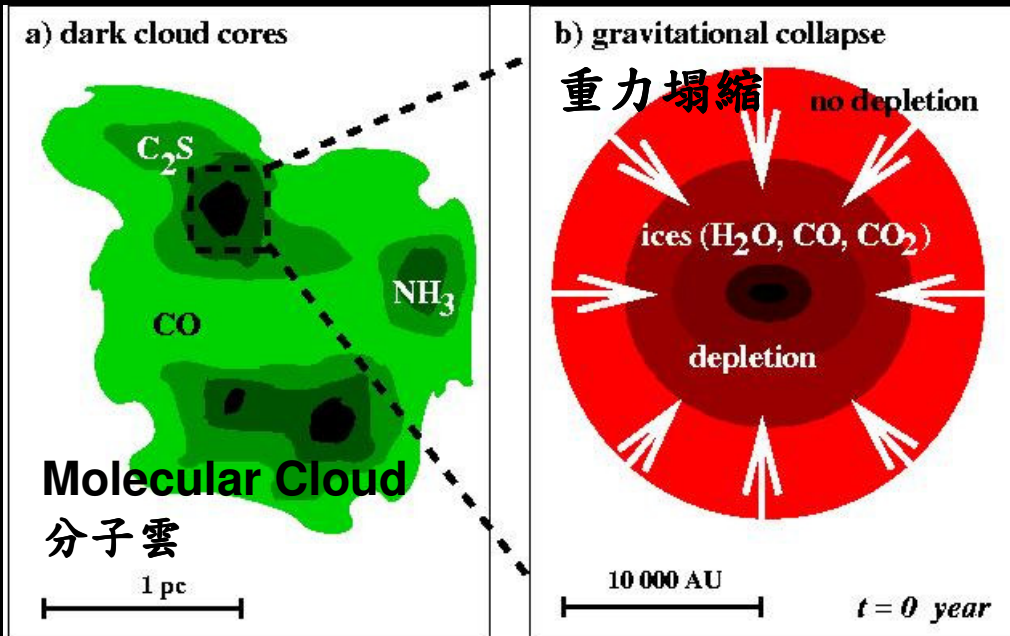


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Scenario of Star & Planet formation



A star is made of gas → star formation needs gas → a star forming region is full of dense gas and dust (原料).

Formation of the Solar System

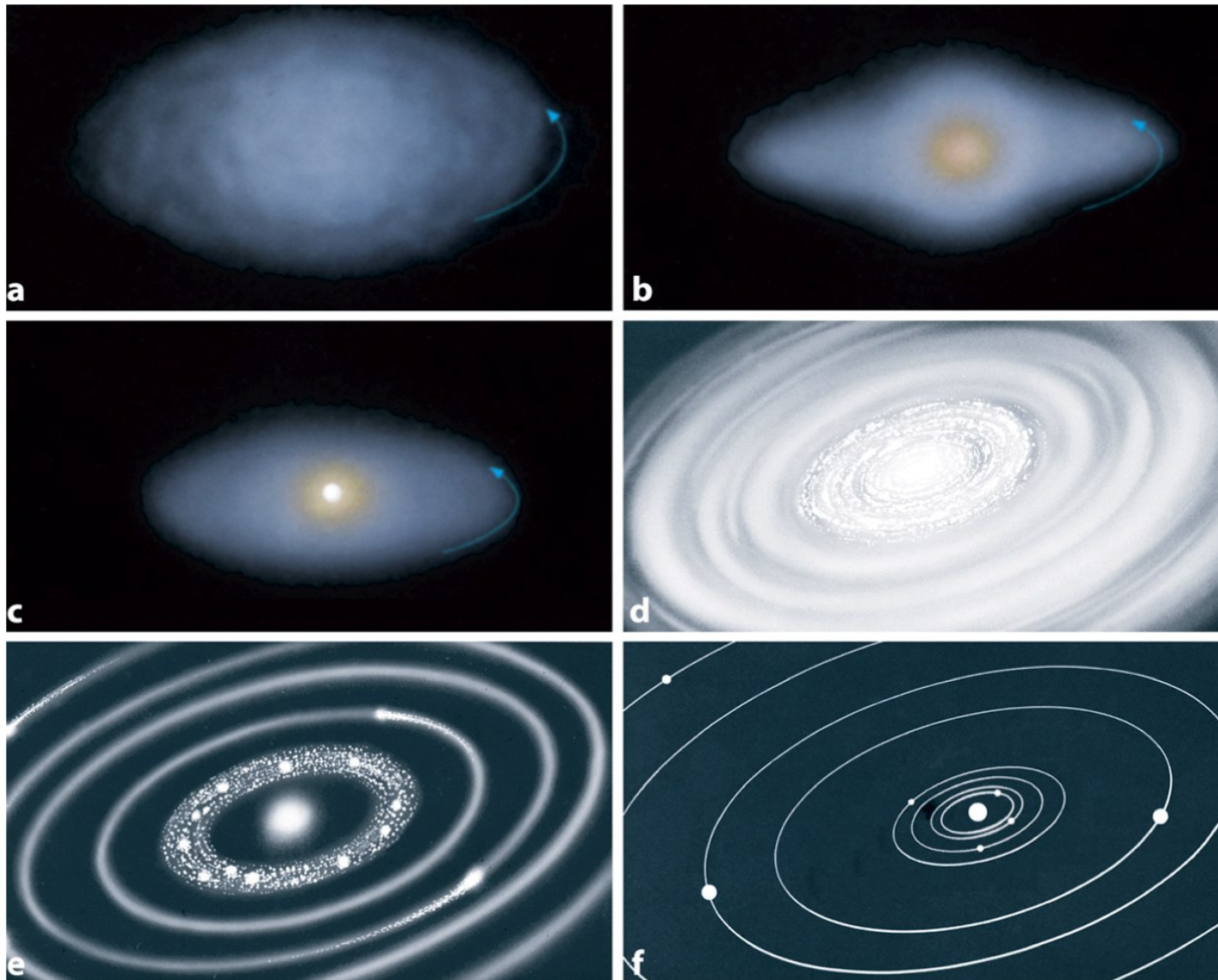


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Formation of planetesimals

Within the disk that surrounds the protosun, solid grains collide and clump together into planetesimals.

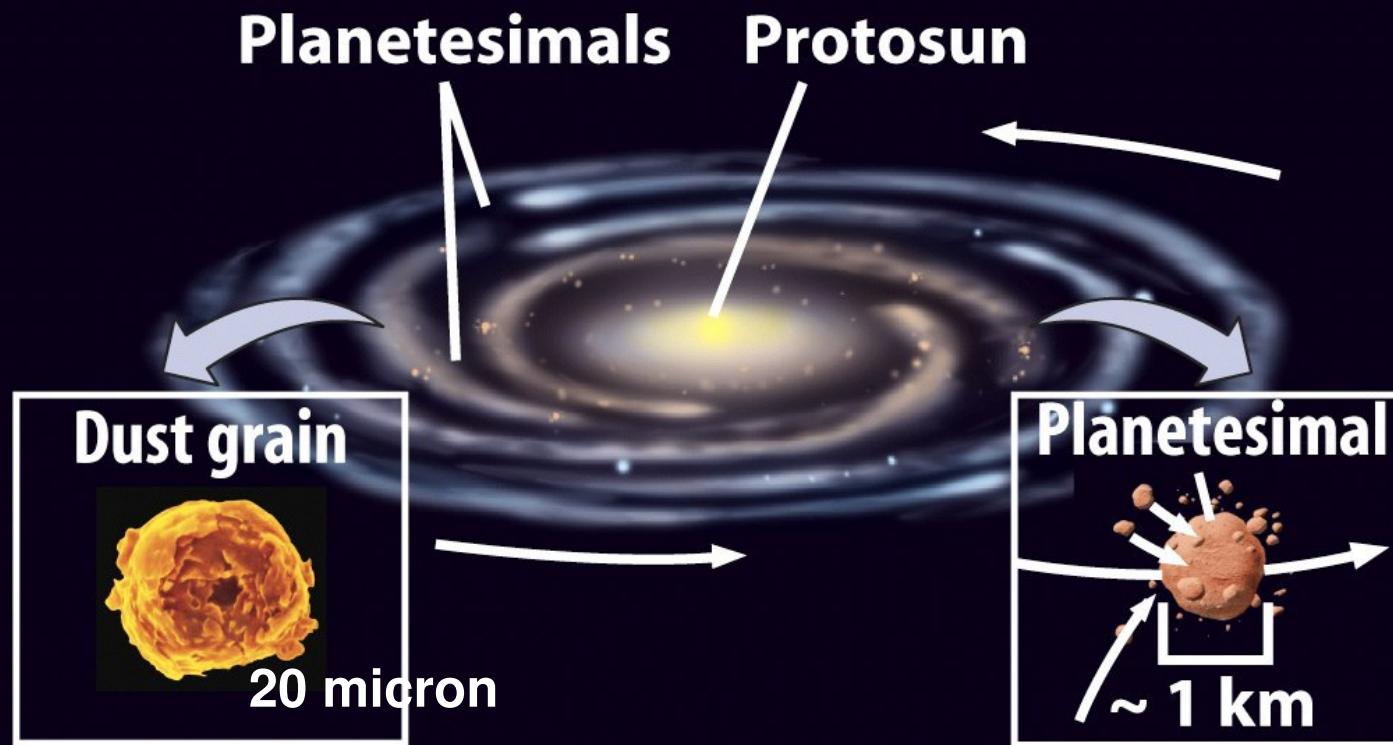


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Formation of planets

The terrestrial planets built up by collisions and by the accretion of planetesimals by gravitational attraction. The Jovian planets formed by gas accretion.

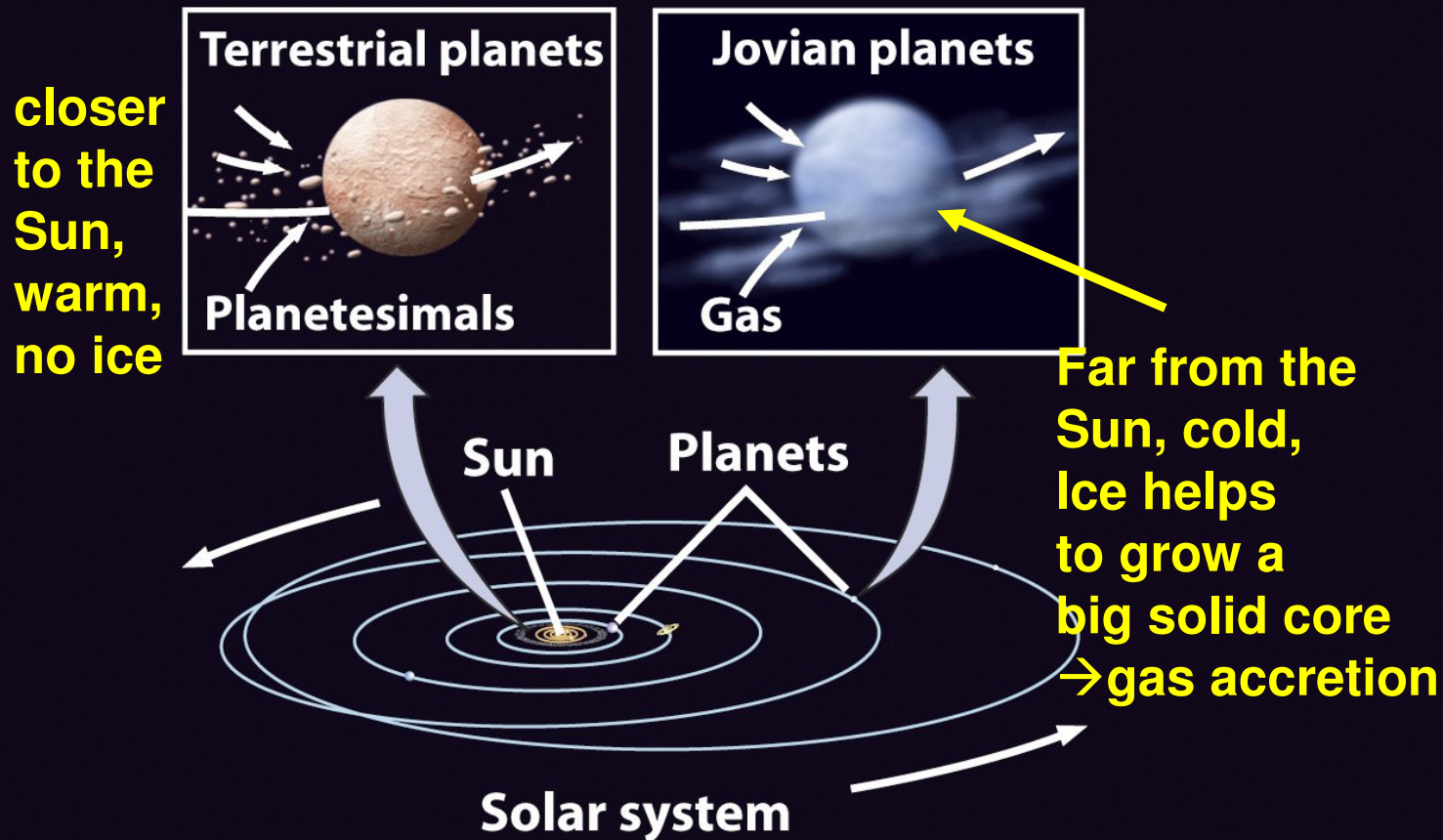
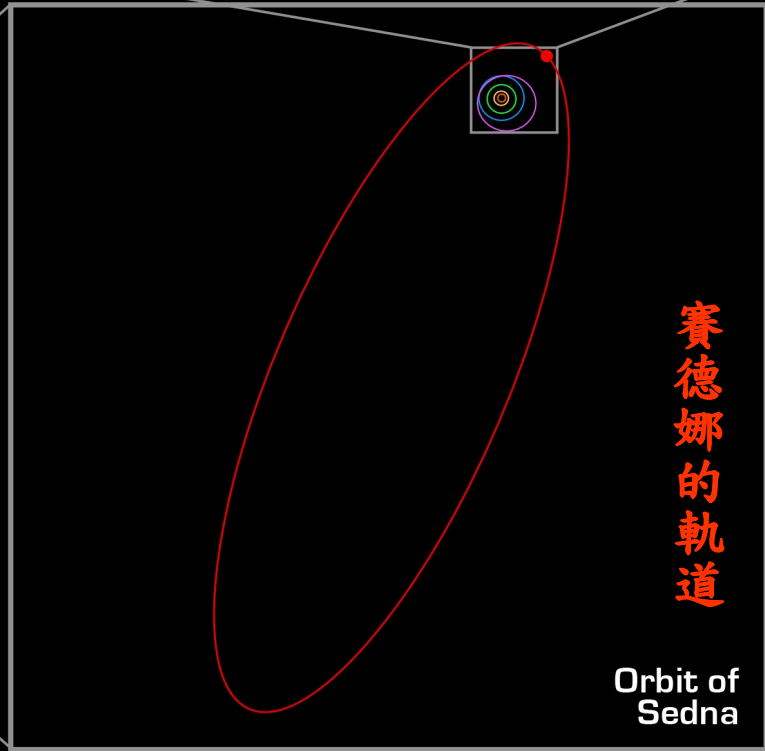
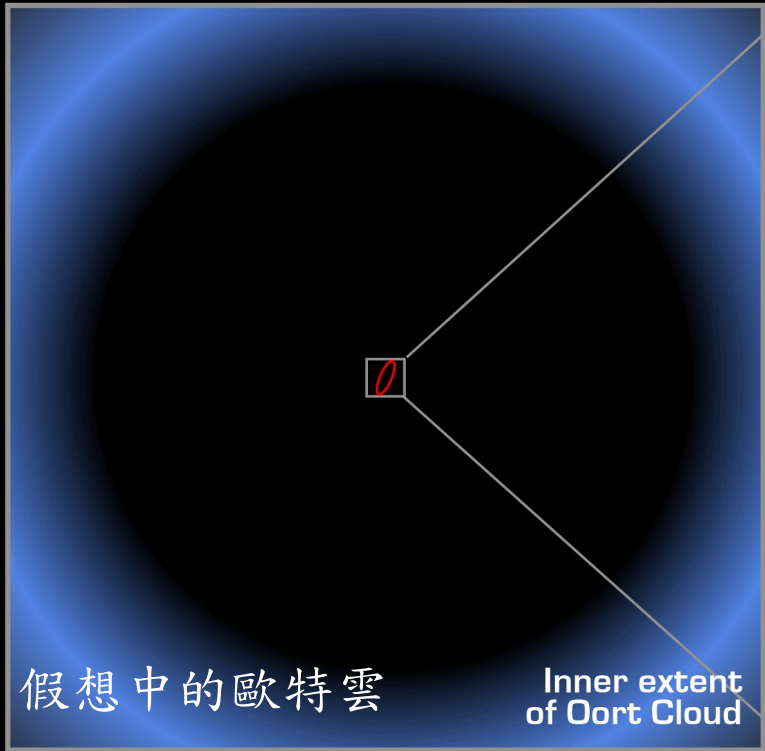
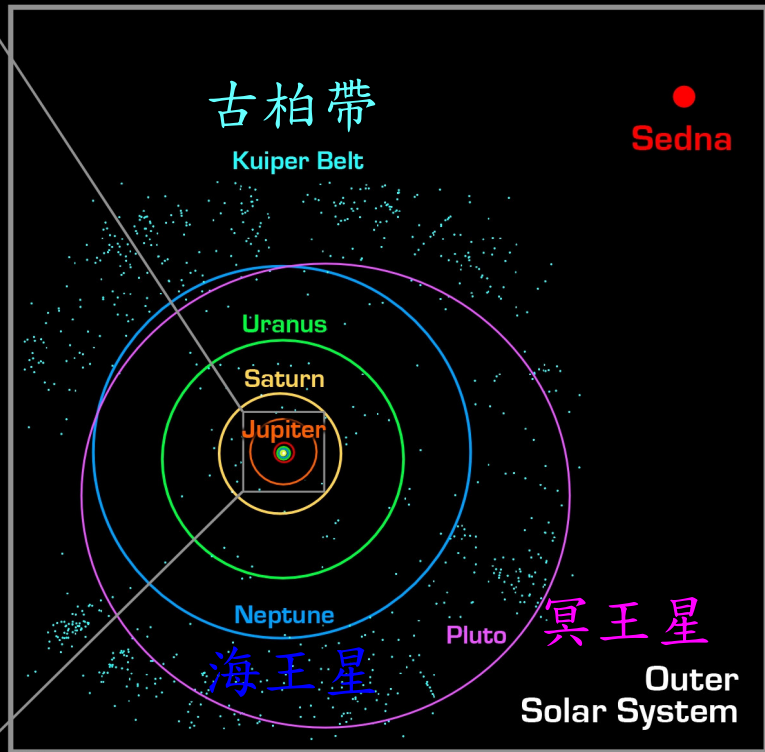
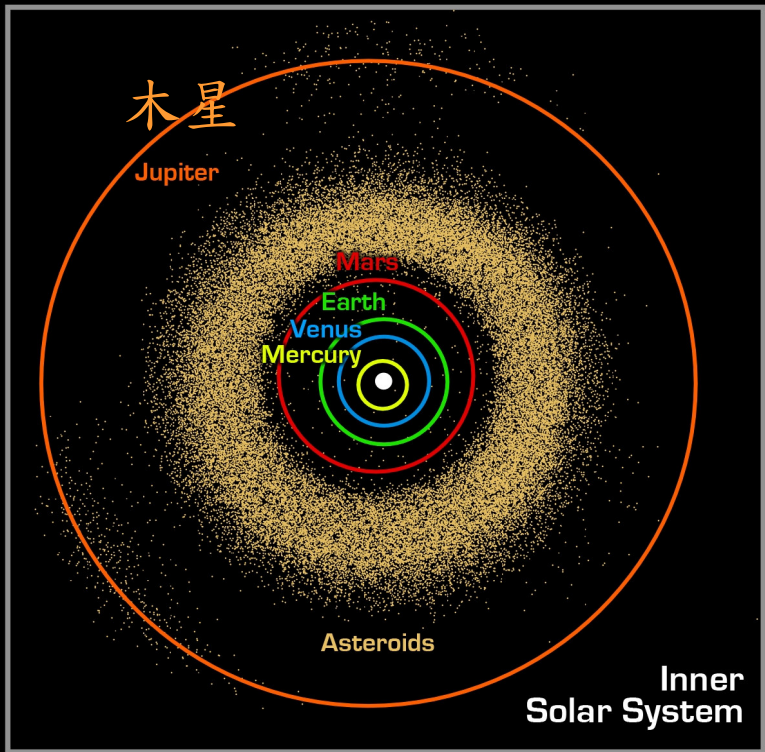


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太陽系

The Solar System

圖片來源：
<http://www.gps.caltech.edu/~mbrown/>

Pluto: the 9th planet?

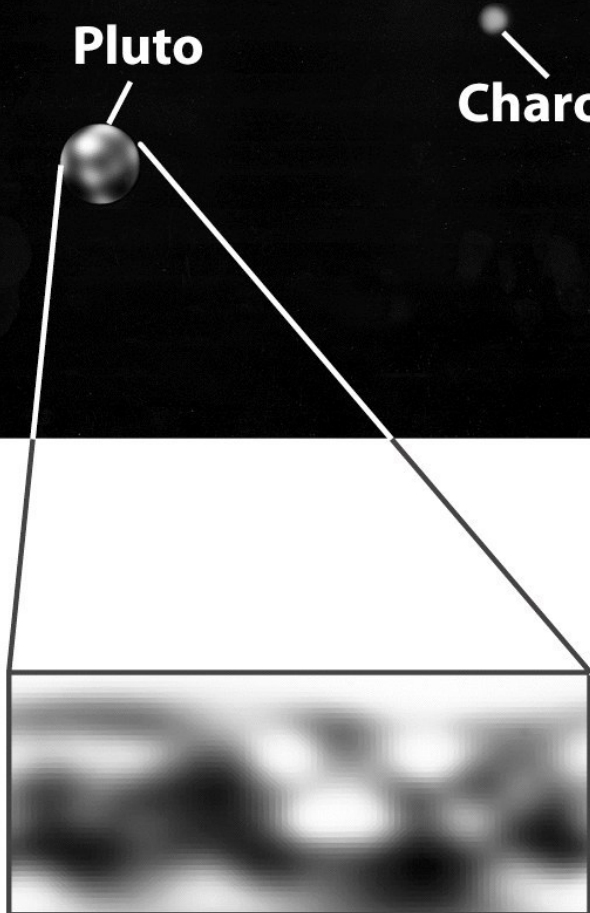
by Hubble Space telescope

Distance to Sun

39.5 AU

Pluto

Charon



1 AU (astronomical unit)
=earth-to-Sun distance

Eris
(satellite:
Dysnomia)

2003 UB₃₁₃
~2,600 km

Sedna
~1,600 km

Charon
1,250 km

Quaoar
~1,250 km

Pluto
2,250 km

Moon
3,476 km

© 2005 Sky & Telescope

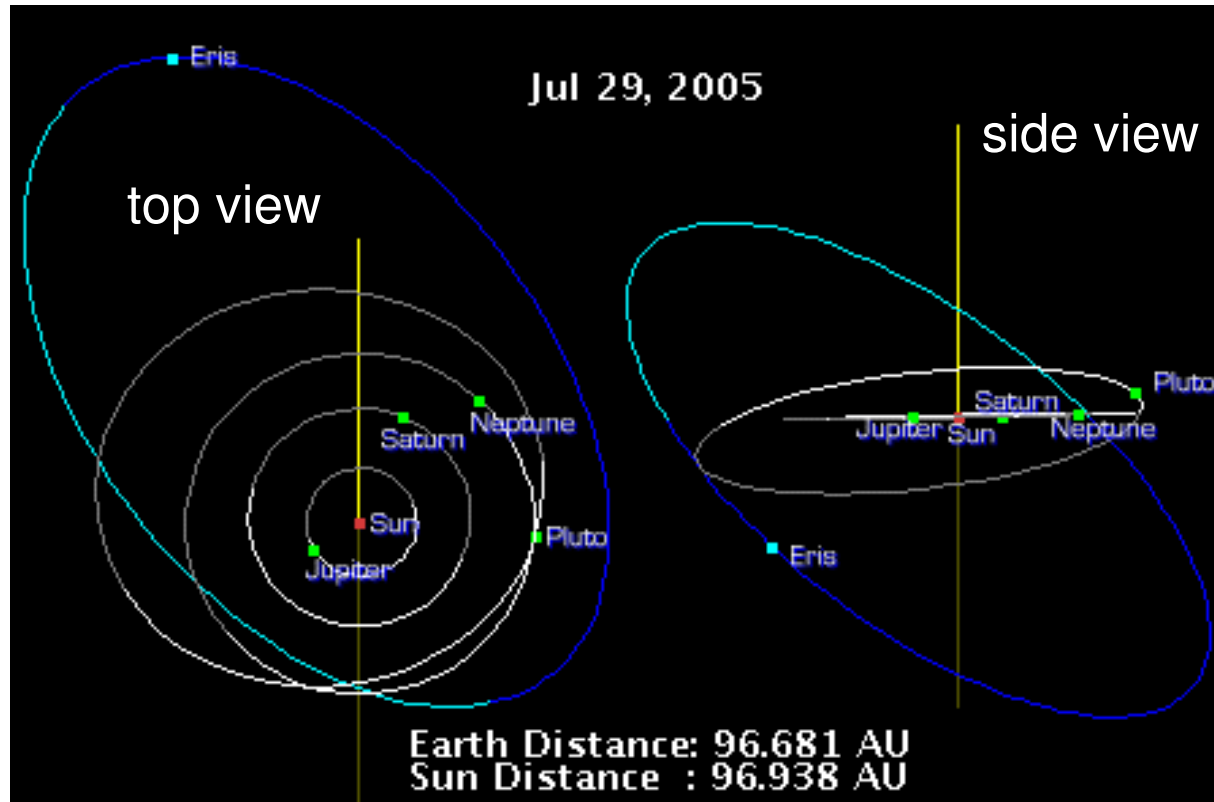
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辜品高：星星・月亮・太陽

8

Orbit of Eris



<http://bobthealien.co.uk/eris.htm>

New Planet Definition

<http://www.iau2006.org/mirror/www.iau.org/iau0603/index.html>

- ◆ orbit a star (upper mass limit: brown dwarf 棕矮星, no nuclear fusion?)
- ◆ massive enough to keep a round shape (lower mass limit: about 5×10^{23} g or 半徑 > 400 km ?)
- ◆ has cleared the neighborhood around its orbit.

Pluto, Eris, and Ceres (穀神星) are now classified as “dwarf planets”. They are the same as “planets” except that firstly they can’t clear the neighborhood & secondly they are not satellites.

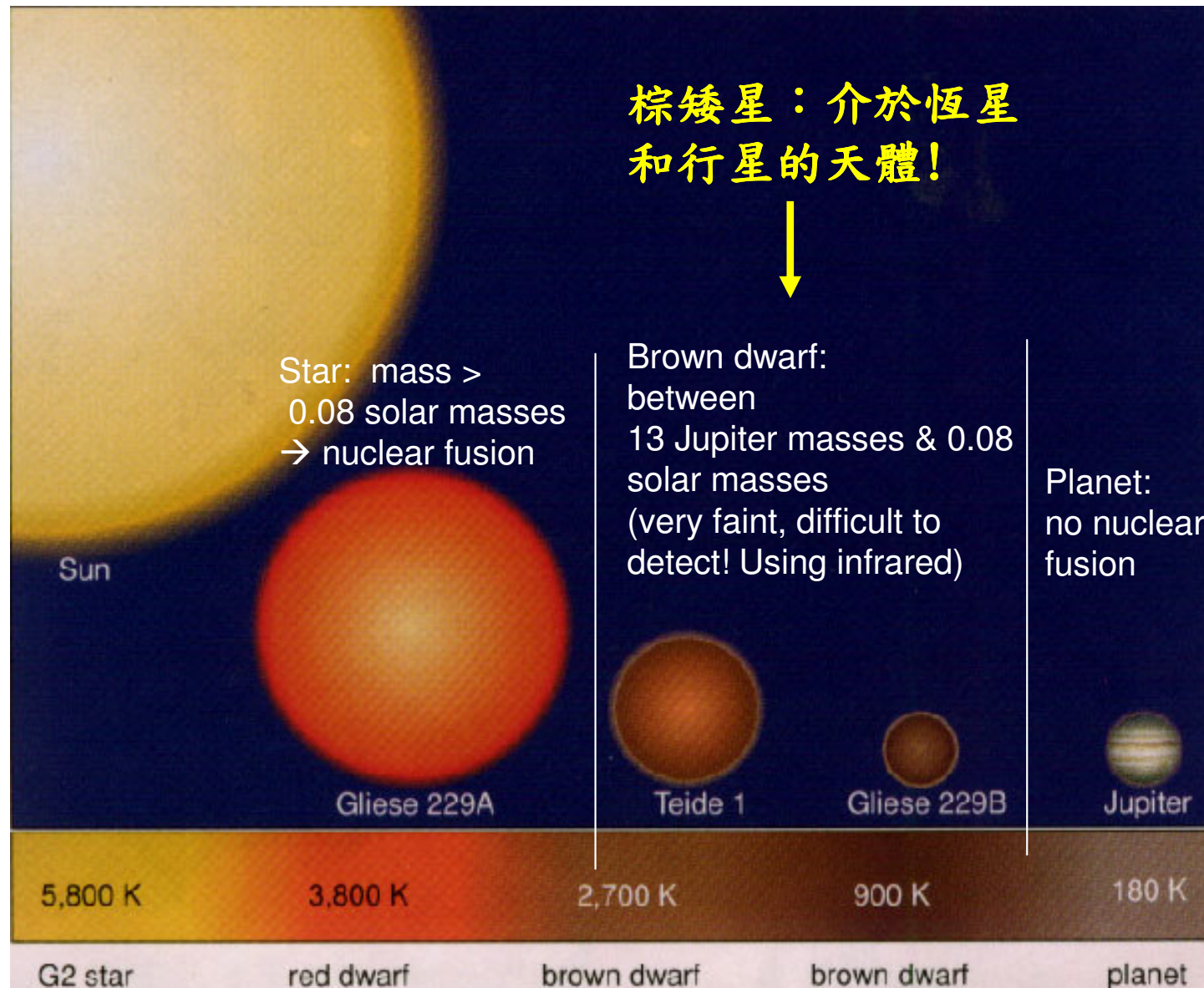
Other Important Issues:

- thermal view: small planets may preserve the raw materials of planet formation
- formation view: Ceres & KBOs may fail to form an “original planet” in the past.
- twin view: what about if the planet mass is comparable to its satellite mass?

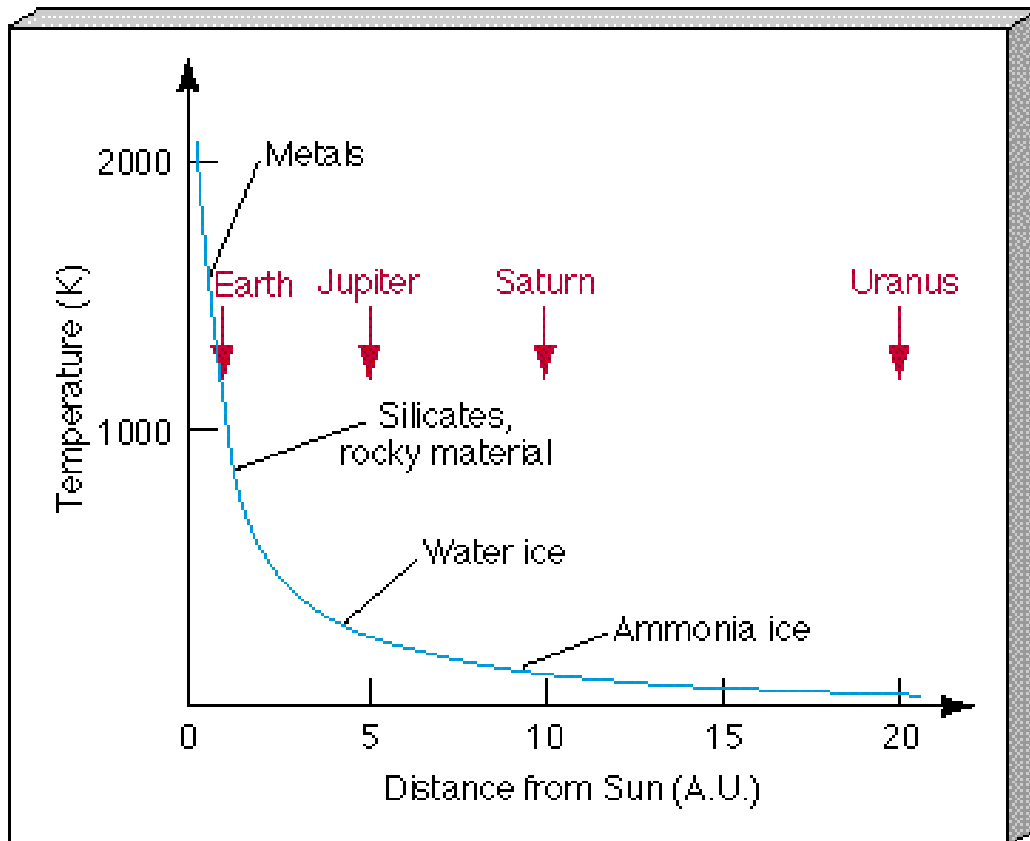
note

In the outskirts of the Solar System, Pluto is accompanied with a group of comet-like bodies at the similar orbital radii. This group is called the Kuiper Belt. The discovery of Eris and the continuous size distribution among largest Kuiper Belt objects suggest that we need to revise the definition of a “planet”. After the IAU (International Astronomical Union) general meeting held this August, astronomers have voted to re-define a “planet”. According to the new definition, now we have only 8 planets. Pluto and Eris are dwarf planets. The number of dwarf planets will increase.

Ultra-low mass stars & Brown dwarfs



Condensation sequence (凝結序列)



273K=0°C

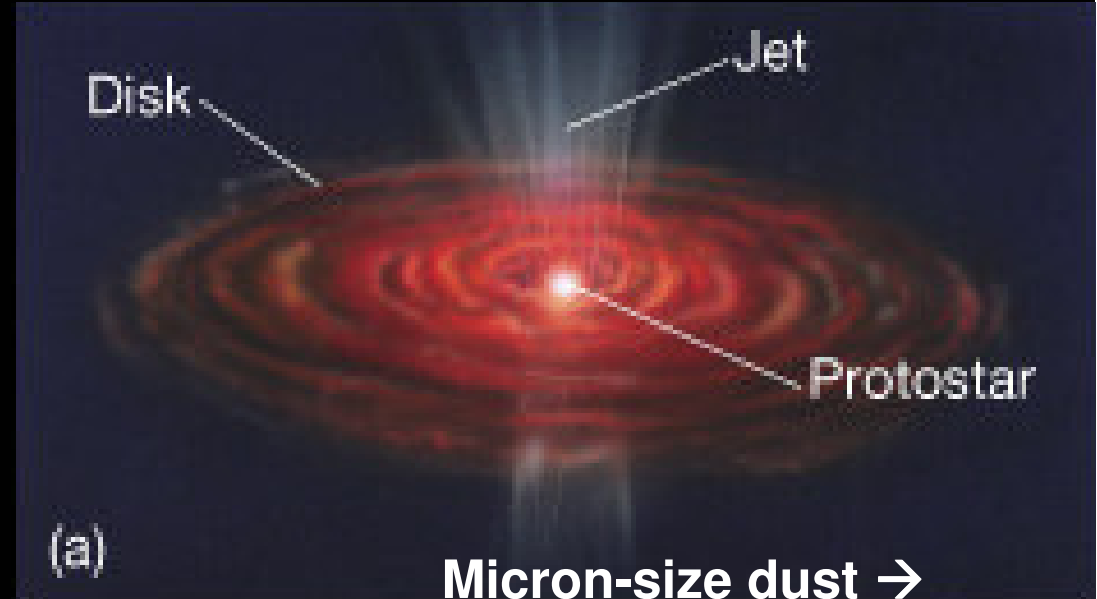
Table 19-3 The Condensation Sequence

Temperature (K)	Condensate	Planet (Estimated temperature of Formation; K)
1500	Metal oxides	Mercury (1400)
1300	Metallic iron and nickel	
1200	Silicates	
1000	Feldspars	Venus (900)
680	Troilite (FeS)	Earth (600) Mars (450)
175	H ₂ O ice	Jovian (175)
150	Ammonia-water ice	
120	Methane-water ice	
65	Argon-neon ice	Pluto (65)

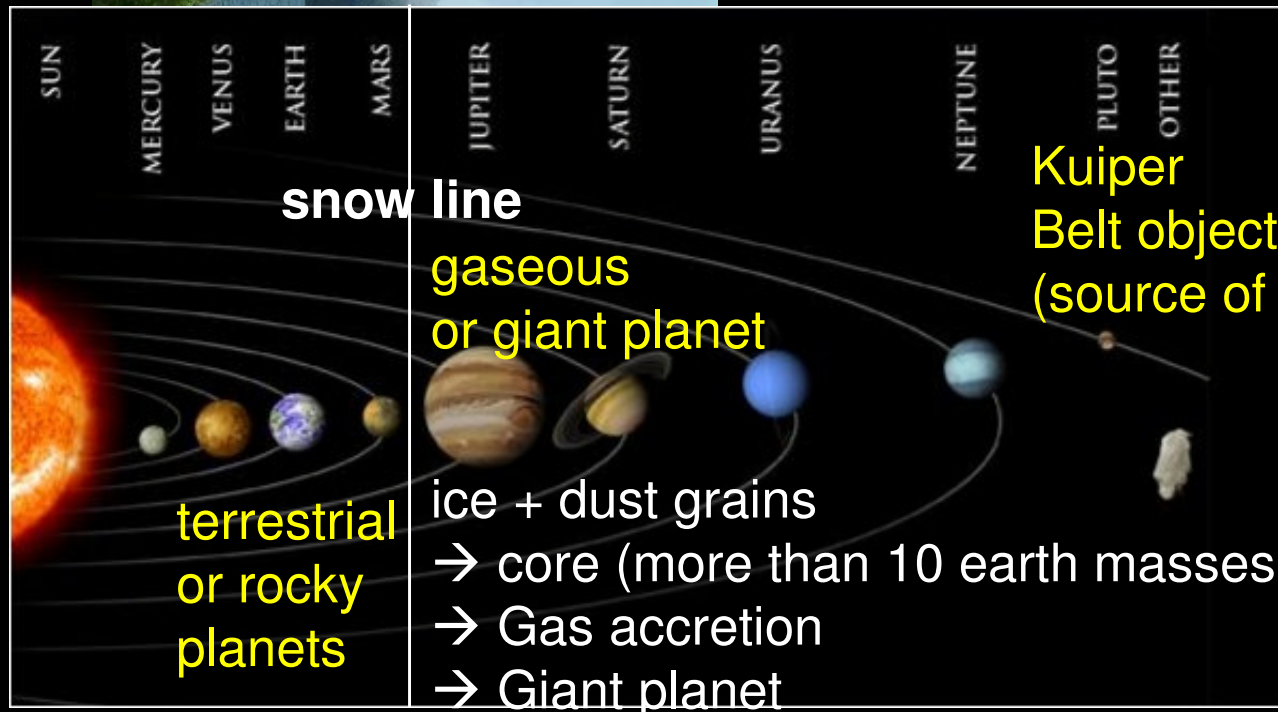
http://ircamera.as.arizona.edu/astr_250/Lectures/Lec_21sml.htm
<http://inverse.astro.uwo.ca/ast21/slides20/slide2.html>

Planets form in a proto-stellar disk around a protostar.

Snow Line on a proto-stellar disk



Micron-size dust →
planetesimals (comets
or asteroids) →
planets



Comet (彗星): dirty snowball

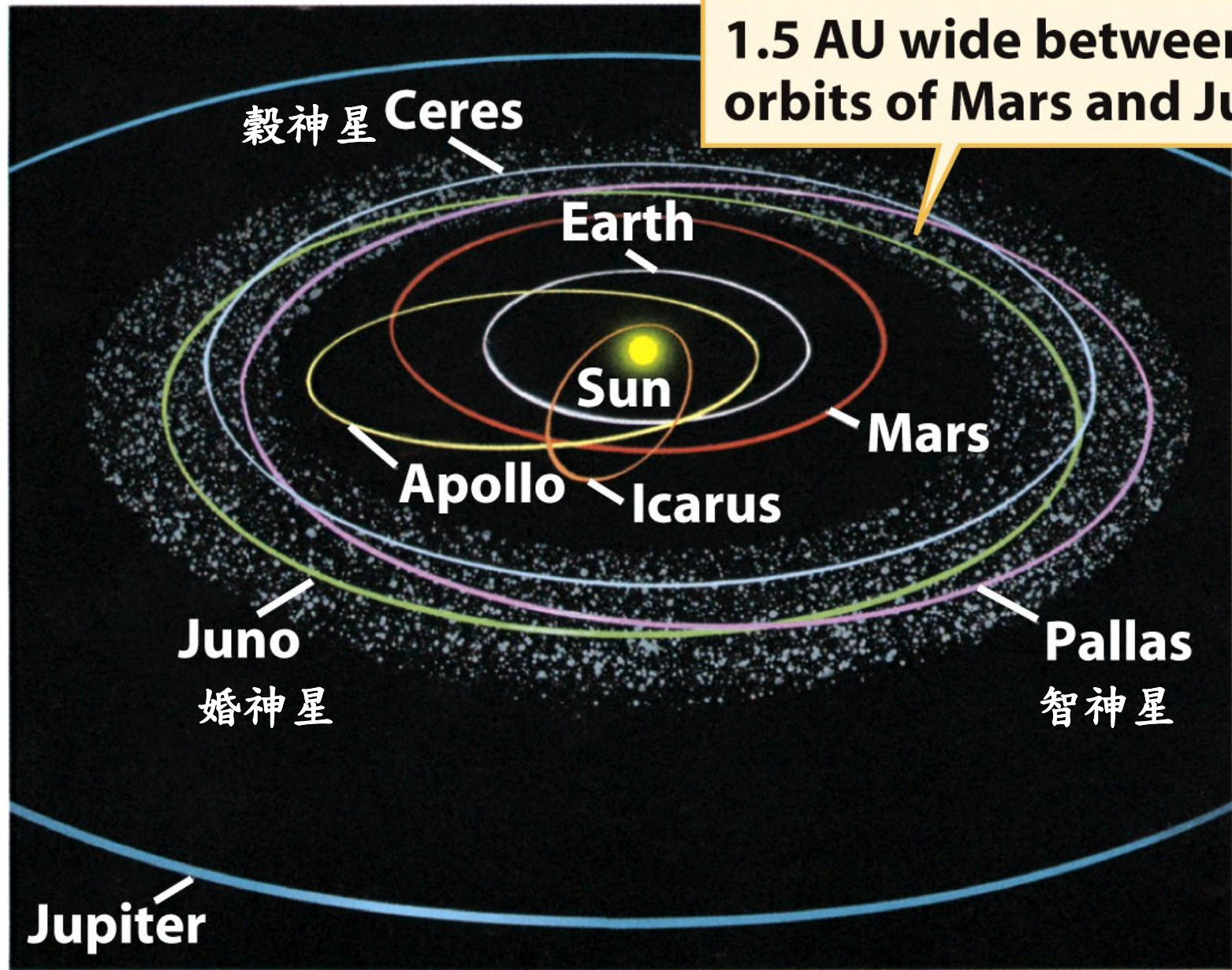


**Fred Whipple
proposed comets
are dirty snowballs
In 1950.**

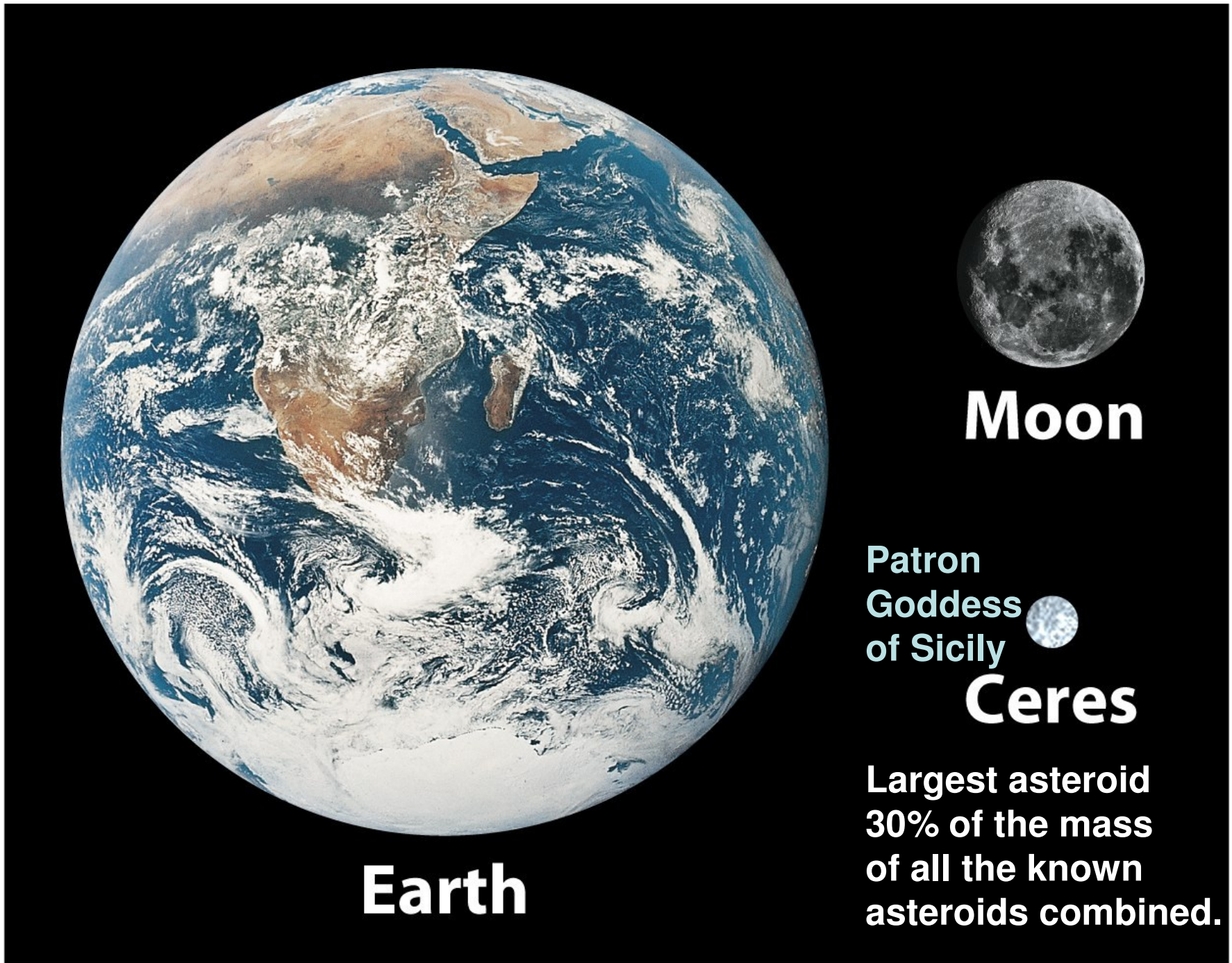
<http://www.spitzer.caltech.edu/Media/releases/ssc2005-18/ssc2005-18b.shtm>

Asteroid: minor planet
小行星 (c.f. planetesimal)

Most asteroids orbit the Sun in a belt about 1.5 AU wide between the orbits of Mars and Jupiter.



2006/11/ Figure 9-2a
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Earth



Moon

Patron
Goddess
of Sicily



Ceres

**Largest asteroid
30% of the mass
of all the known
asteroids combined.**

An Asteroid has a “moon”

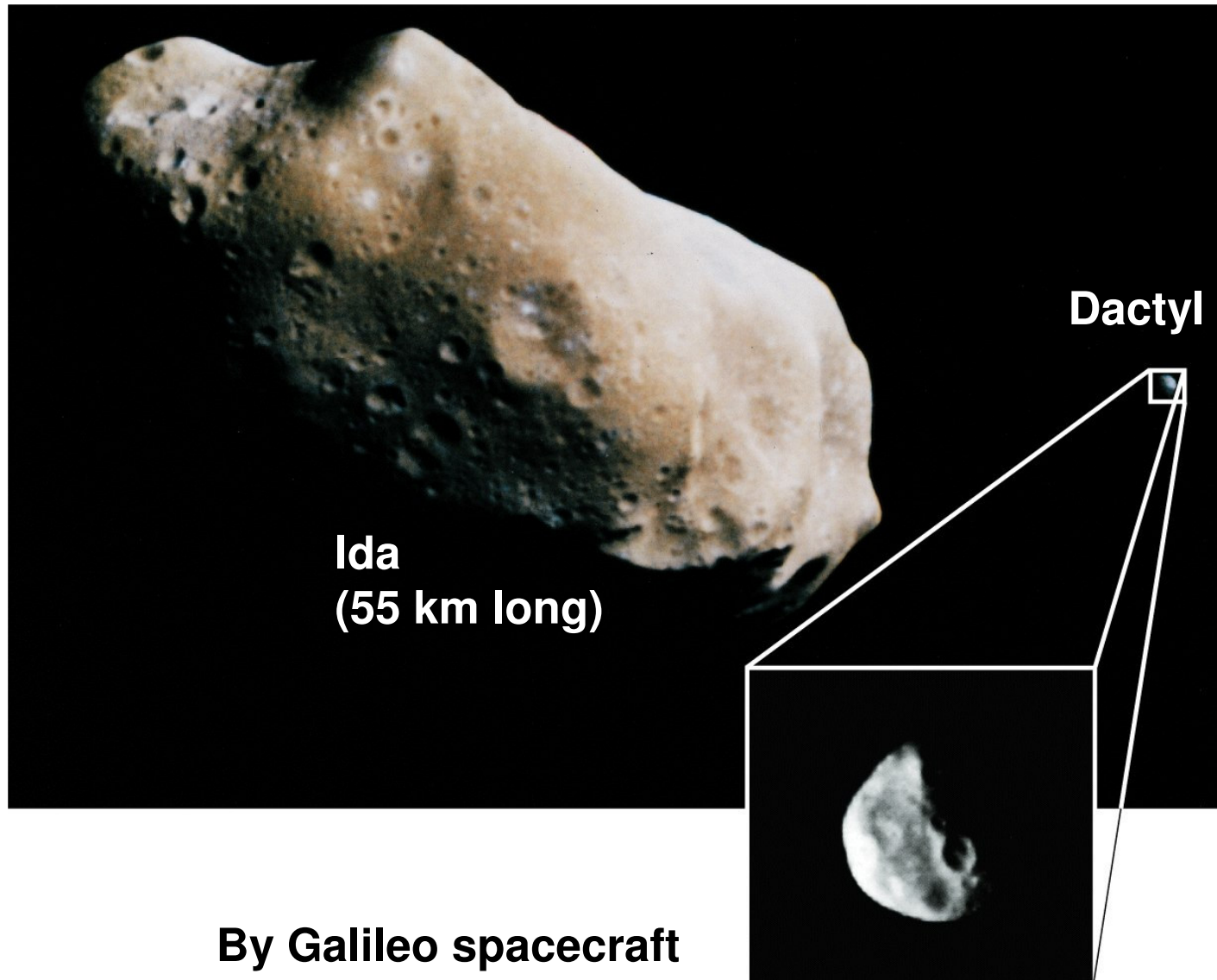
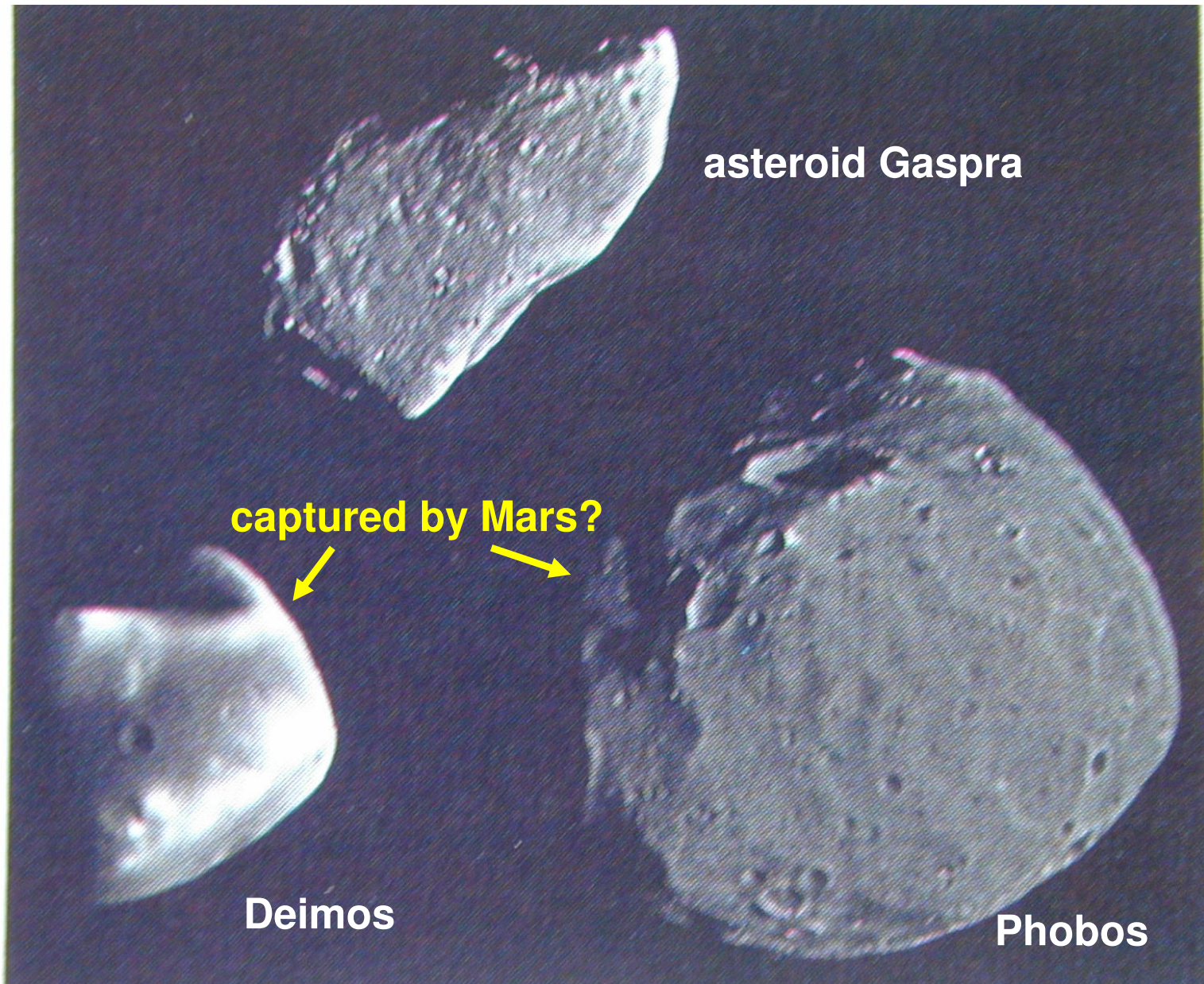


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Asteroid & Mars's Moons



Name an asteroid

- Normally, the discoverer of an asteroid has the first right to name it after someone else, but recent automated telescopic surveys gave the IAU (International Astronomical Union) a surplus of asteroids with nothing but prosaic catalog designations. Anyone can petition the IAU to name an asteroid in someone's honor, but the process often takes years. Also, IAU's Committee for Small Body Nomenclature has specific rules for names, including a limit of 16 characters (including punctuation), a requirement that the name be pronounceable (in some language), and no names from pet animals allowed.
- For instance, 3 asteroids were named in Memory of September 11 terrorism victims:
http://www.space.com/news/asteroid_setp11_011010.html

Properties of belt asteroids

- Number: as of 2002, 52224 asteroids confirmed, more than 125000 awaiting confirmation.
- Average separation: 10 million km
- Composition & Albedo: S, C, M types
- Size: irregular shapes. 30 asteroids between 200-300 km in diameter. The number increases with decreasing size. Putting together \rightarrow $\frac{1}{2}$ the diameter of our Moon.
- A failed planet? (recall how a planet formed?)

Composition of asteroids

Consist primarily of rock & metal.

Compared with comets which are composed of rock and ices (dirty snowball).

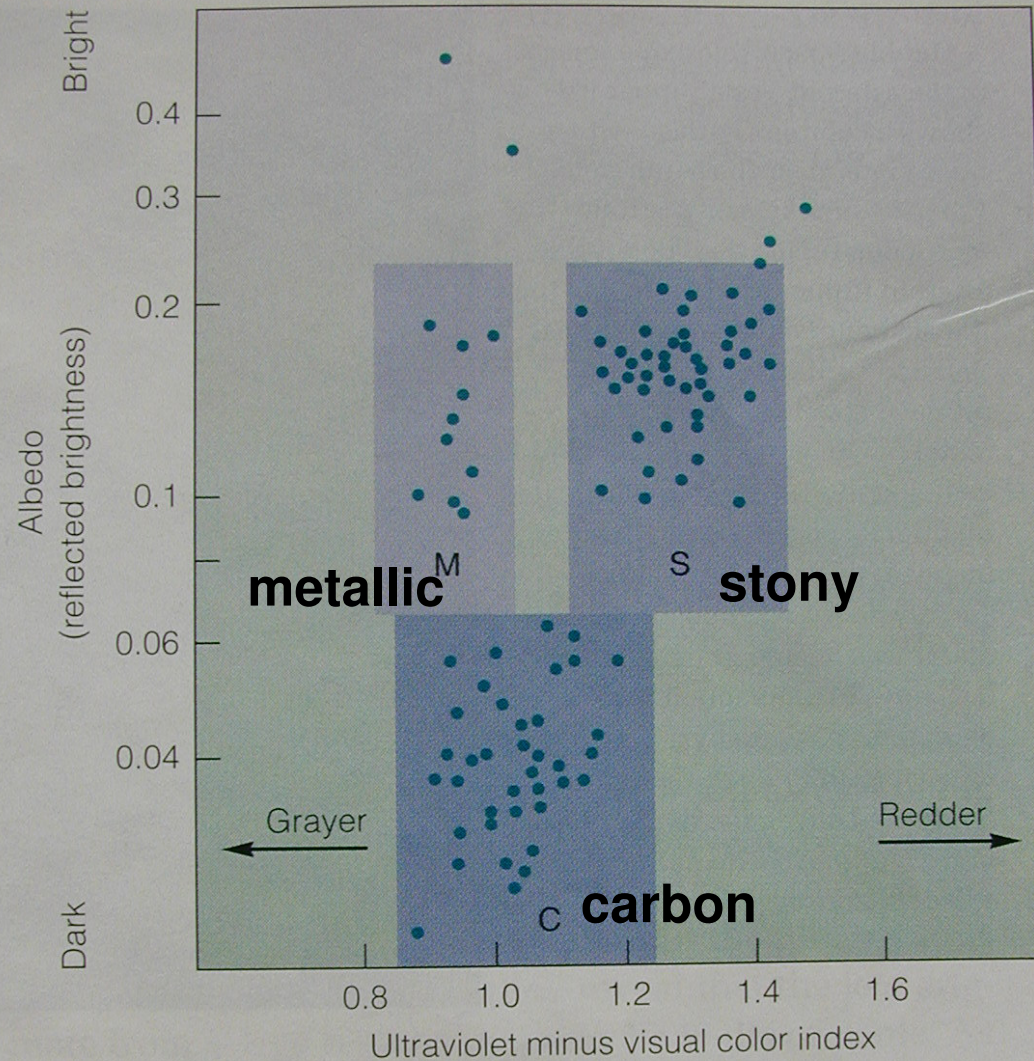


FIGURE 9-6

The three principal types of asteroids. In a diagram of albedo versus color (ultraviolet minus visual), the most reflective asteroids lie near the top, and the reddest at the right. The S types, which are believed to resemble chondrites, are clearly redder than the M types, which seem to be metallic. The darkest asteroids, the C types, are believed to be similar to carbonaceous chondrites. (Diagram adapted from a figure by B. Zellner)

Trojan (特洛伊) asteroids

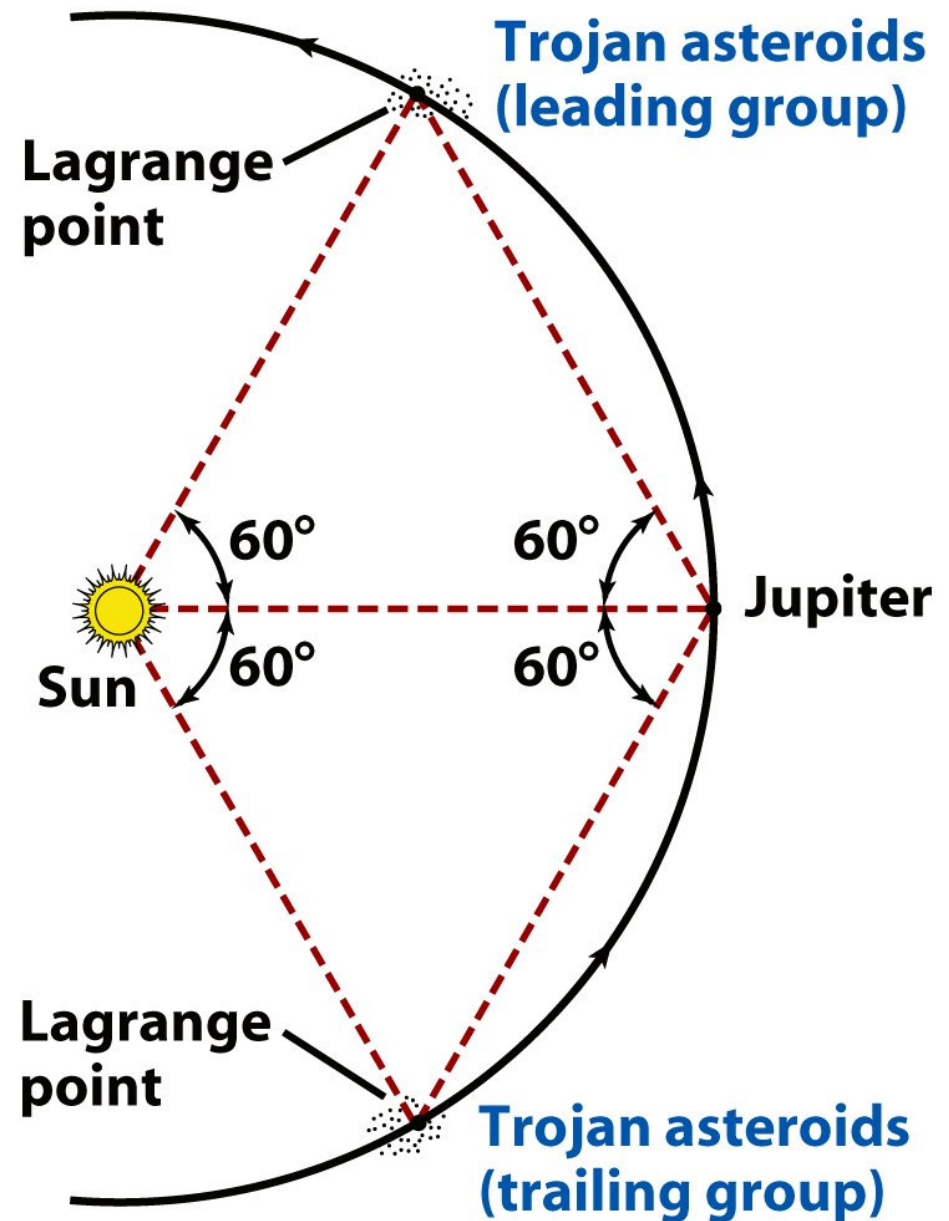


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Landing on an asteroid

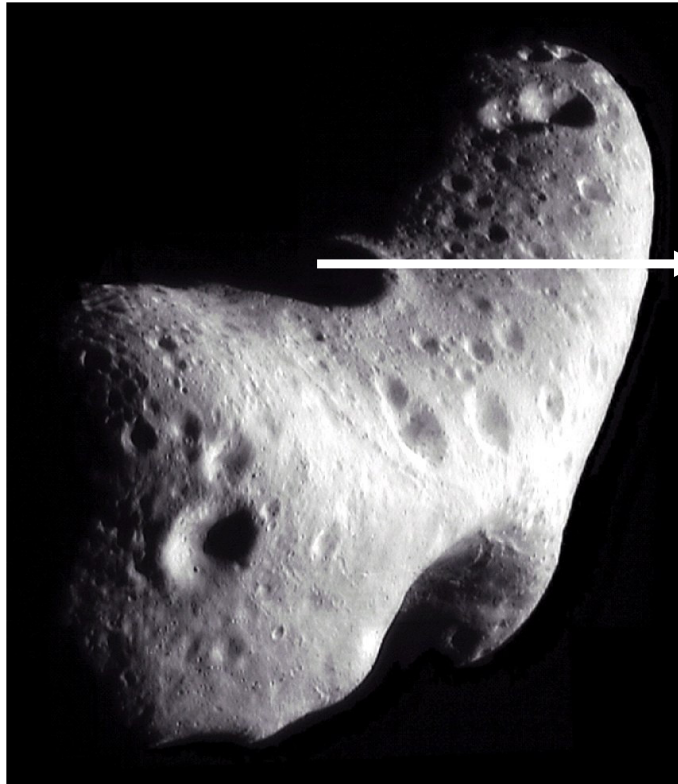


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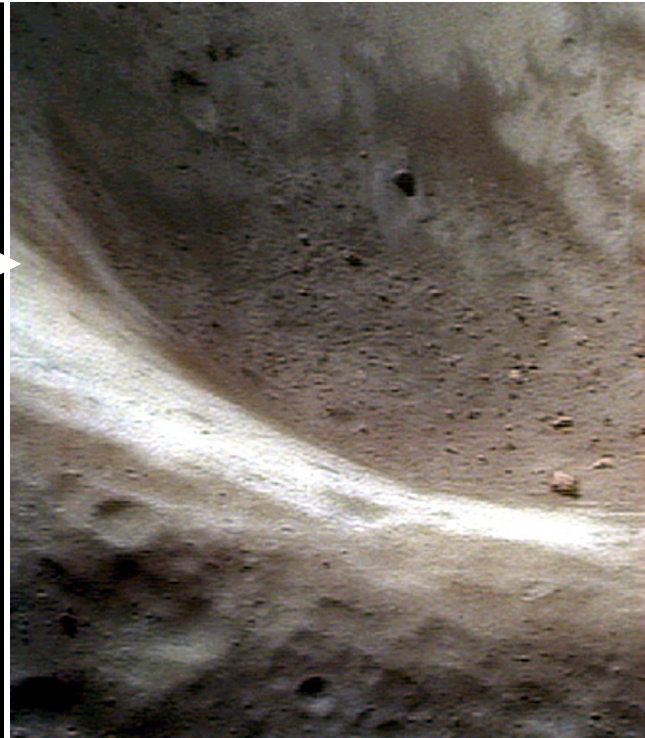


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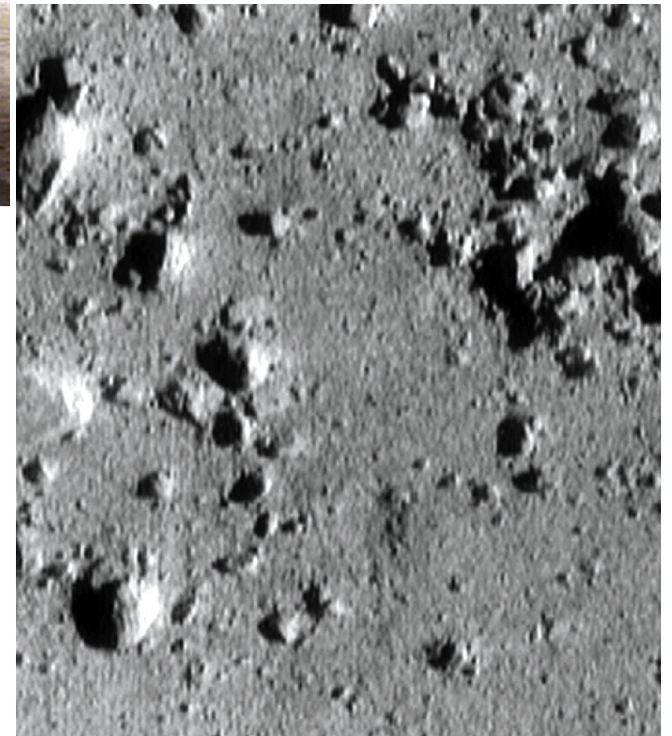


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Eros(愛神): 33x13x13km
5.24 hours/rotation

Near Earth Asteroid Rendezvous Shoemaker
(NEAR)

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辜品高：星星・月亮・太陽

Close encounter

Asteroid 1994 XM1
was less than $\frac{1}{2}$
the distance from
the Earth to the
Moon!

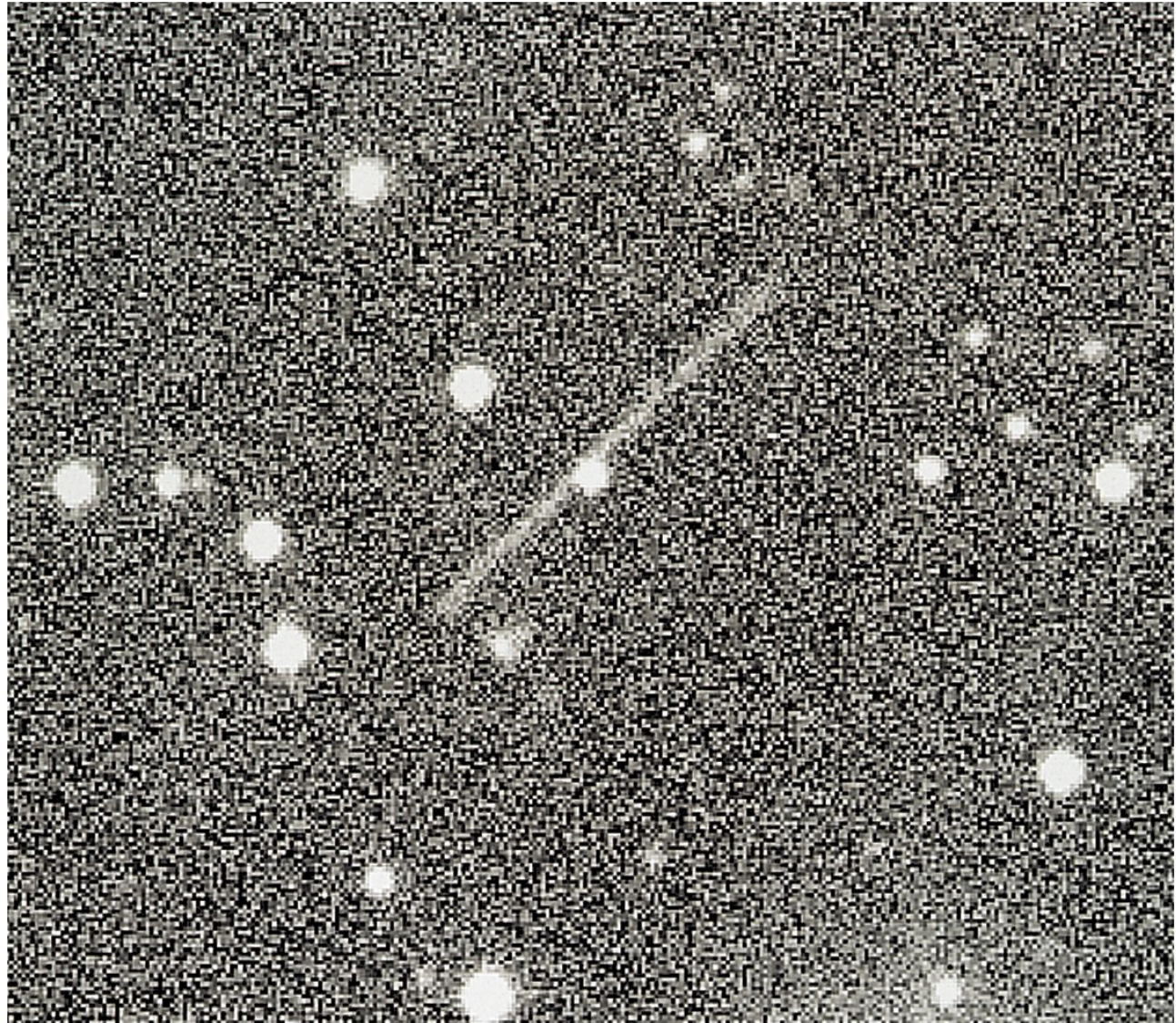


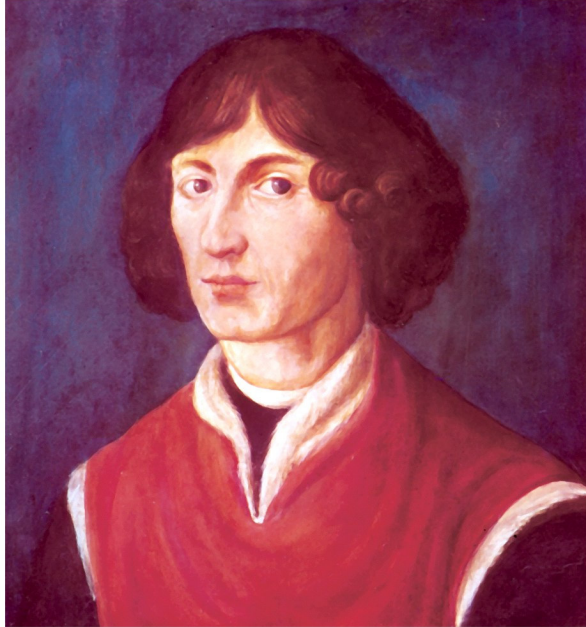
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Near-earth objects (NEOs)

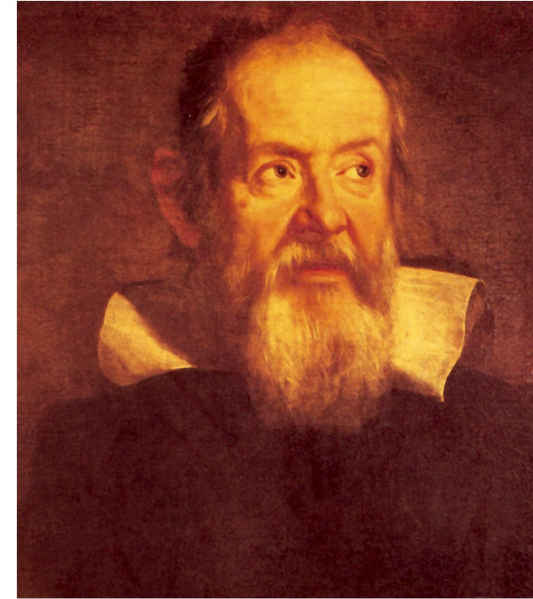
<http://neo.jpl.nasa.gov/>

Group	Description	Definition
NECs	Near-Earth Comets	$q < 1.3$ AU, $P < 200$ years
NEAs	Near-Earth Asteroids	$q < 1.3$ AU
Atens	Earth-crossing NEAs with semi-major axes smaller than Earth's (named after asteroid 2062 Aten).	$a < 1.0$ AU, $Q > 0.983$ AU
Apollos	Earth-crossing NEAs with semi-major axes larger than Earth's (named after asteroid 1862 Apollo).	$a > 1.0$ AU, $q < 1.017$ AU
Amors	Earth-approaching NEAs with orbits exterior to Earth's but interior to Mars' (named after asteroid 1221 Amor).	$a > 1.0$ AU, $1.017 < q < 1.3$ AU
PHAs There are currently known 730 PHAs.	Potentially Hazardous Asteroids: NEAs whose Minimum Orbit Intersection Distance (MOID) with the Earth is 0.05 AU (about 20 times the distance between Earth & Moon) or less and whose <u>absolute magnitude (H)</u> is 22.0 or brighter.	$MOID \leq 0.05$

Astronomers (A.D. 1500-1700)



**Copernicus
(1473-1543)
Heliocentric
theory**



**Galileo
(1564-1642)
First use a
telescope
to watch the
sky**



**Tycho
(1546-1601)
Half helio +
Half geo
&
Kepler
(1571-1630)
3 laws**

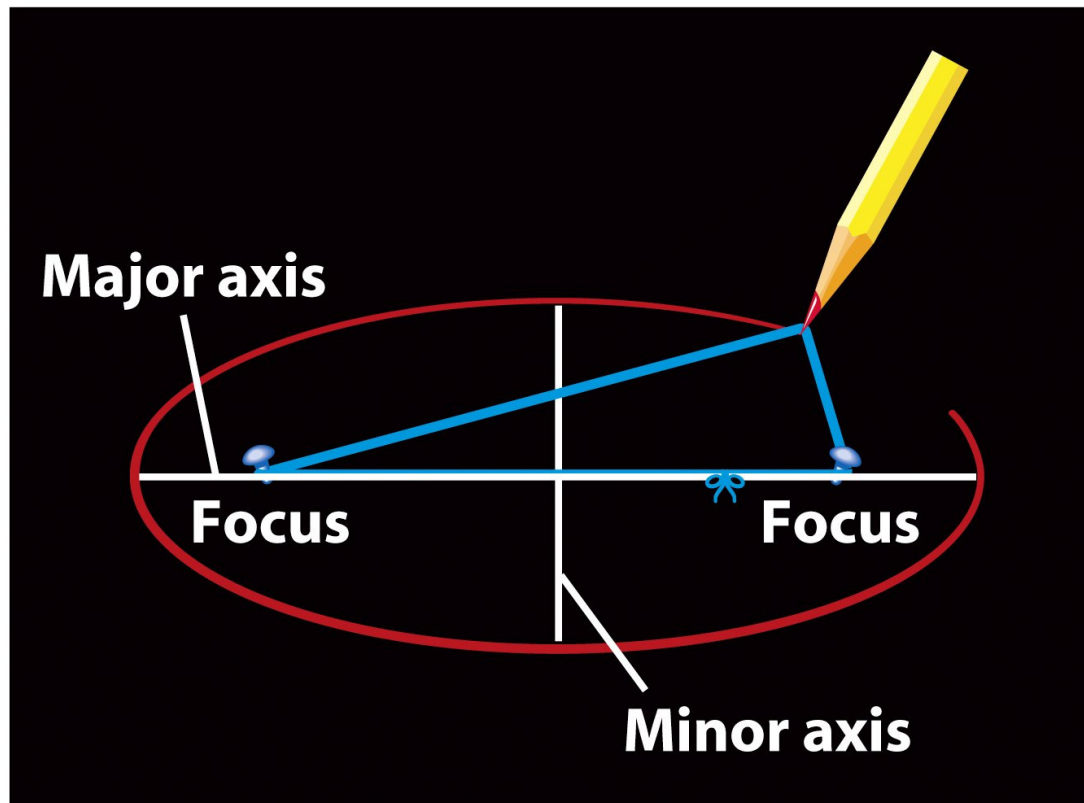
Unnumbered Figure pg 43c
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**Newton
(1642-1727)
Law of
universal
gravitation**

Unnumbered Figure pg 43d
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Ellipse



e 是 eccentricity 離(偏)心率

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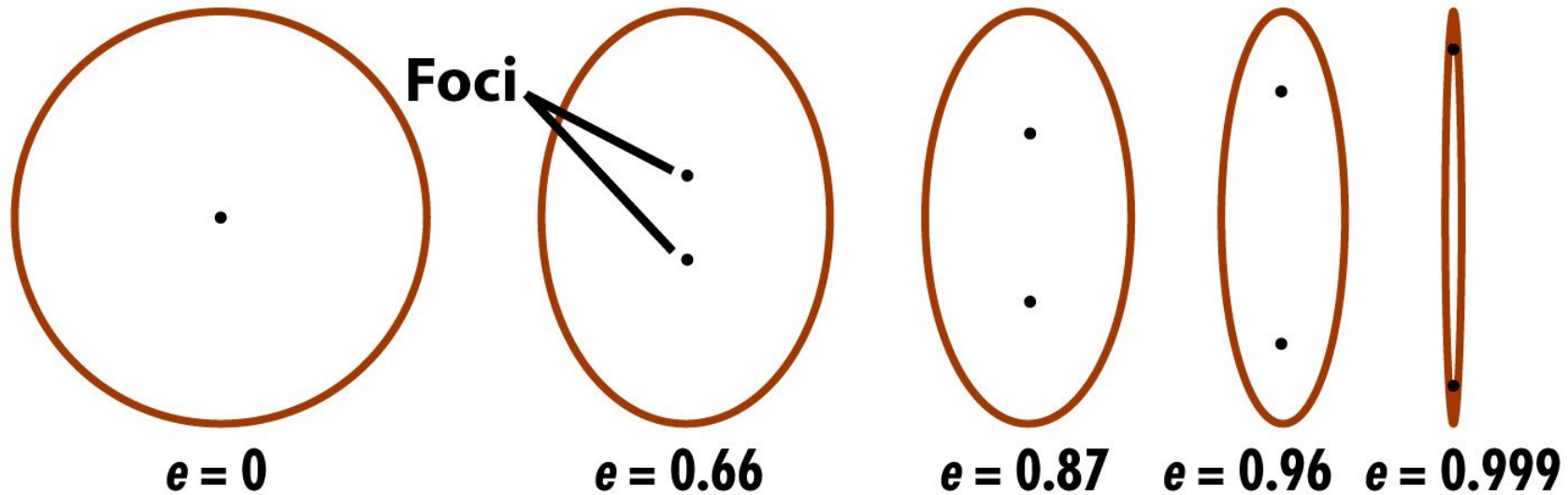


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Kepler's 3 Laws

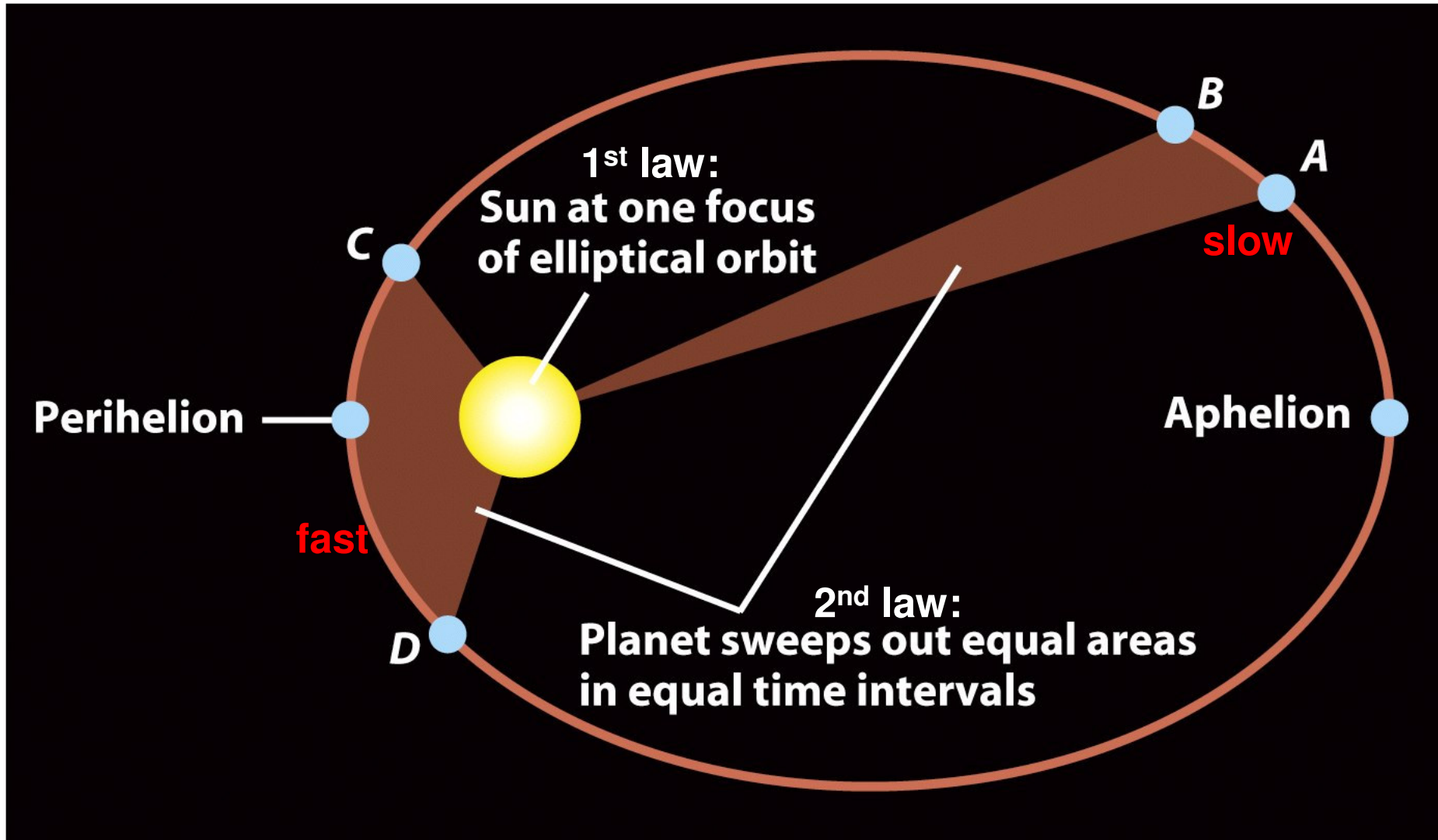


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3rd law: $\text{period}^2 = \text{distance}^3$ the closer, the faster!

(Newton: the smaller the distance, the stronger the gravitational force)

Newton & Gravitation

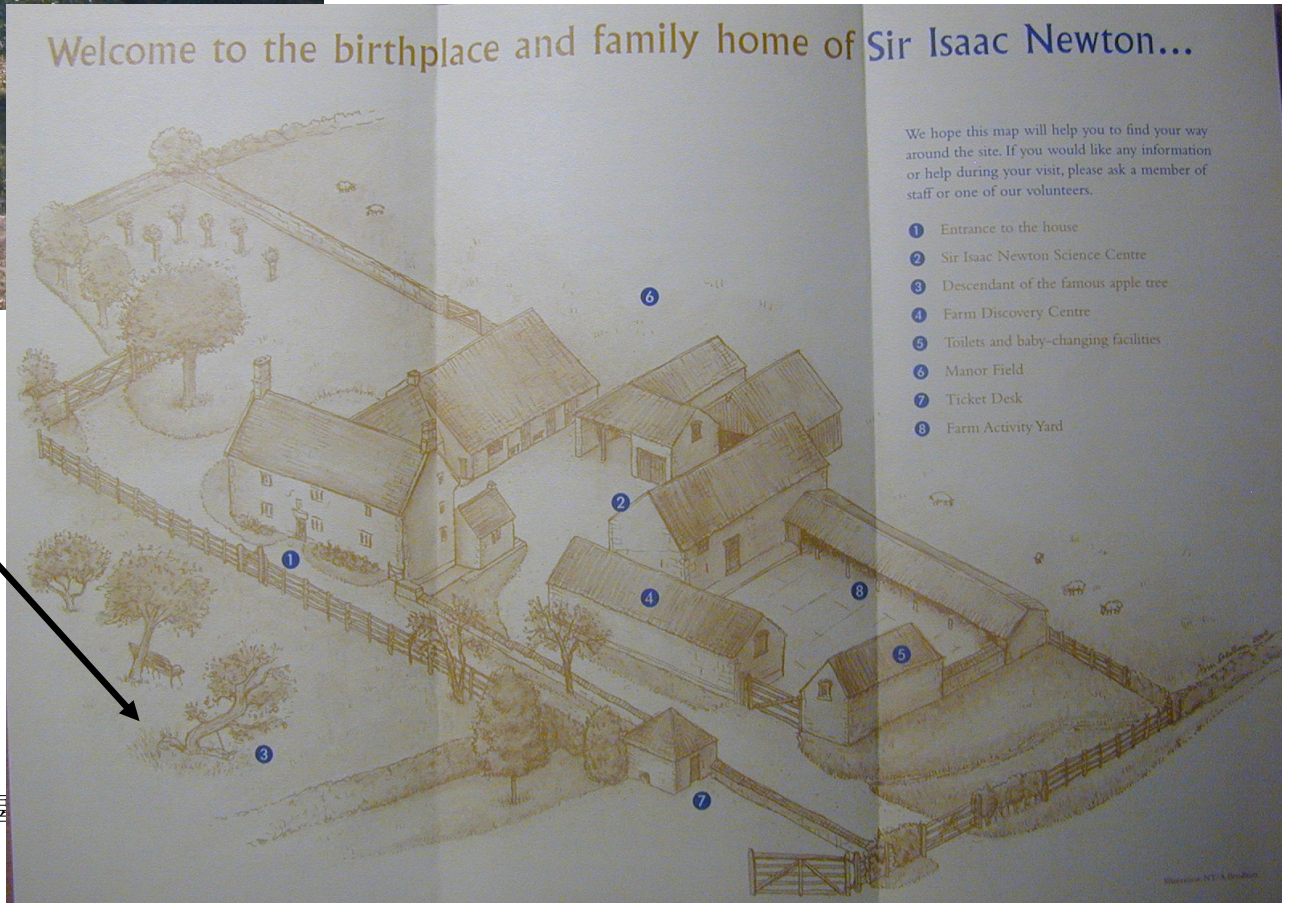
Woolsthorpe



Welcome to the birthplace and family home of Sir Isaac Newton...

We hope this map will help you to find your way around the site. If you would like any information or help during your visit, please ask a member of staff or one of our volunteers.

- 1 Entrance to the house
- 2 Sir Isaac Newton Science Centre
- 3 Descendant of the famous apple tree
- 4 Farm Discovery Centre
- 5 Toilets and baby-changing facilities
- 6 Manor Field
- 7 Ticket Desk
- 8 Farm Activity Yard



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幸

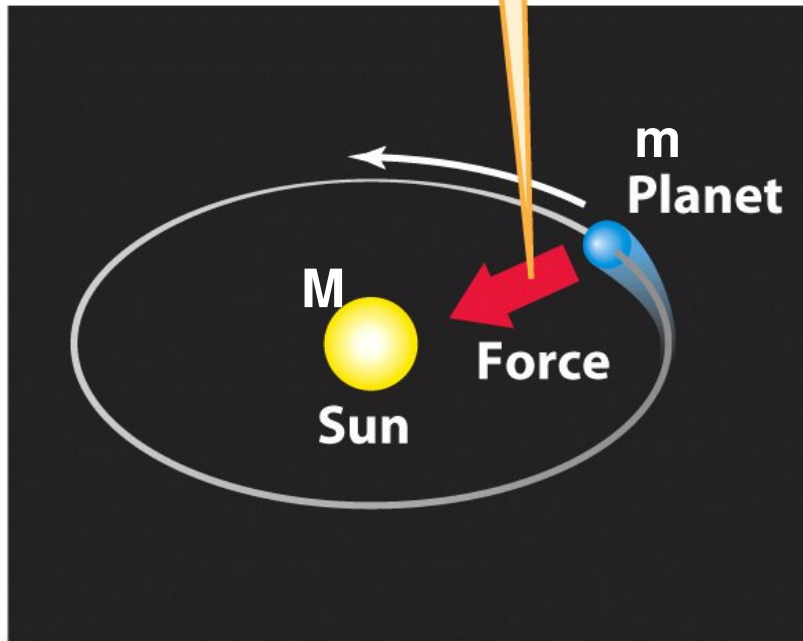
Newton's law of gravitation

$$F_{gravity} \propto \frac{Mm}{r^2} \rightarrow \text{Kepler's laws}$$

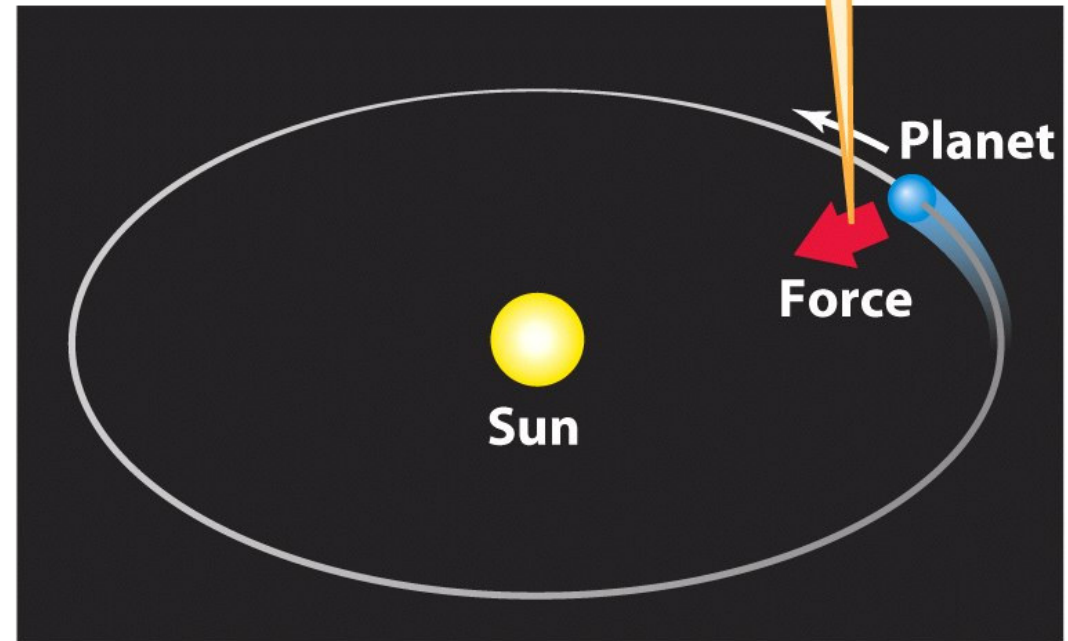
怎麼不會相撞？ Ans:離心力

To make a planet move at a high speed in a small orbit requires a strong gravitational force.

To make the same planet move at a low speed in a larger orbit requires only a weak gravitational force.



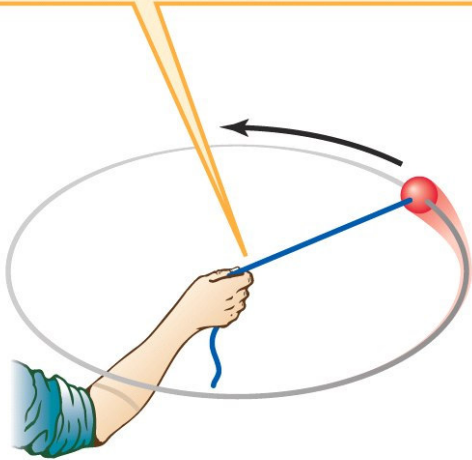
(c)



(d)

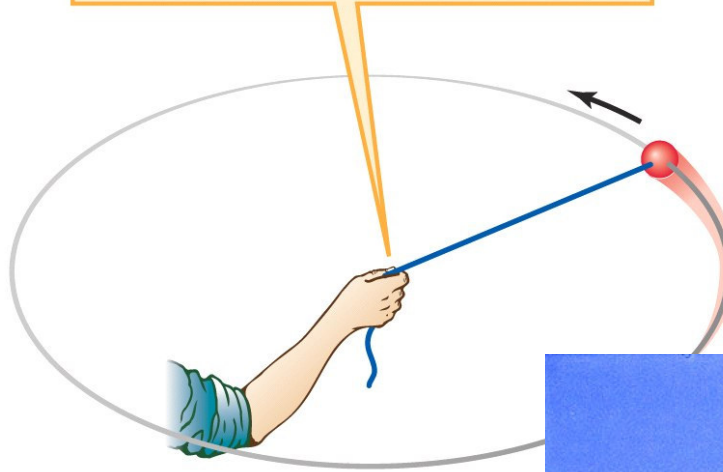
Centrifugal force (離心力)

To make a ball move at a high speed in a small circle requires a strong pull.



(a)

To make the same ball move at a low speed in a large circle requires only a weak pull.



(b)



summary

- 為什麼行星都朝同一個方向以及大略同一個平面做公轉？
- 太陽系除了行星外還有什麼物質？
- 行星形成是由小而大逐漸組合成，還是由大而小破裂成？
- 為什麼說小行星帶可能是早期一羣無法形成行星的小天體？
- 什麼是太陽系的雪線？它對行星的形成和組成有何影響？
- 什麼是特洛伊(Trojan)小行星？
- 什麼是near earth objects?人類有注意它們的行蹤嗎(怕怕)?
- 小行星(asteroid)和慧星(comet)的組成有何不同？
- 為什麼慧星的尾巴與太陽的位置相反？
- 為什麼說古柏帶(Kuiper Belt)是短週期慧星的來緣？
- 為什麼天文學家臆測有歐特雲(Oort Cloud)的存在？
- Eris是第十顆行星嗎？
- 為什麼大型的天體(行星或矮行星)是球形的？
- 有介在恆星和行星之間的天體嗎？
- 什麼是台美掩星計畫？
- 什麼造成流星雨？
- 什麼是牛頓的重力理論？
- 為什麼越接近太陽，行星公轉得越快？